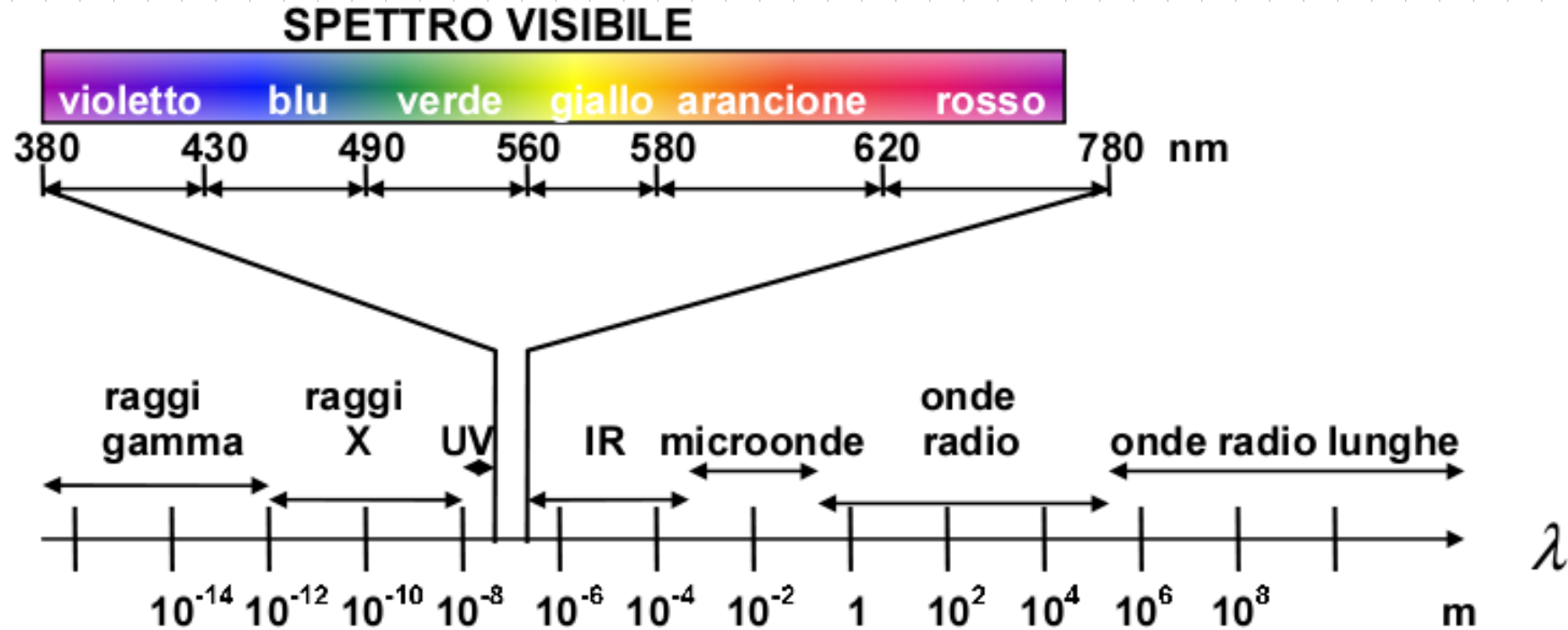
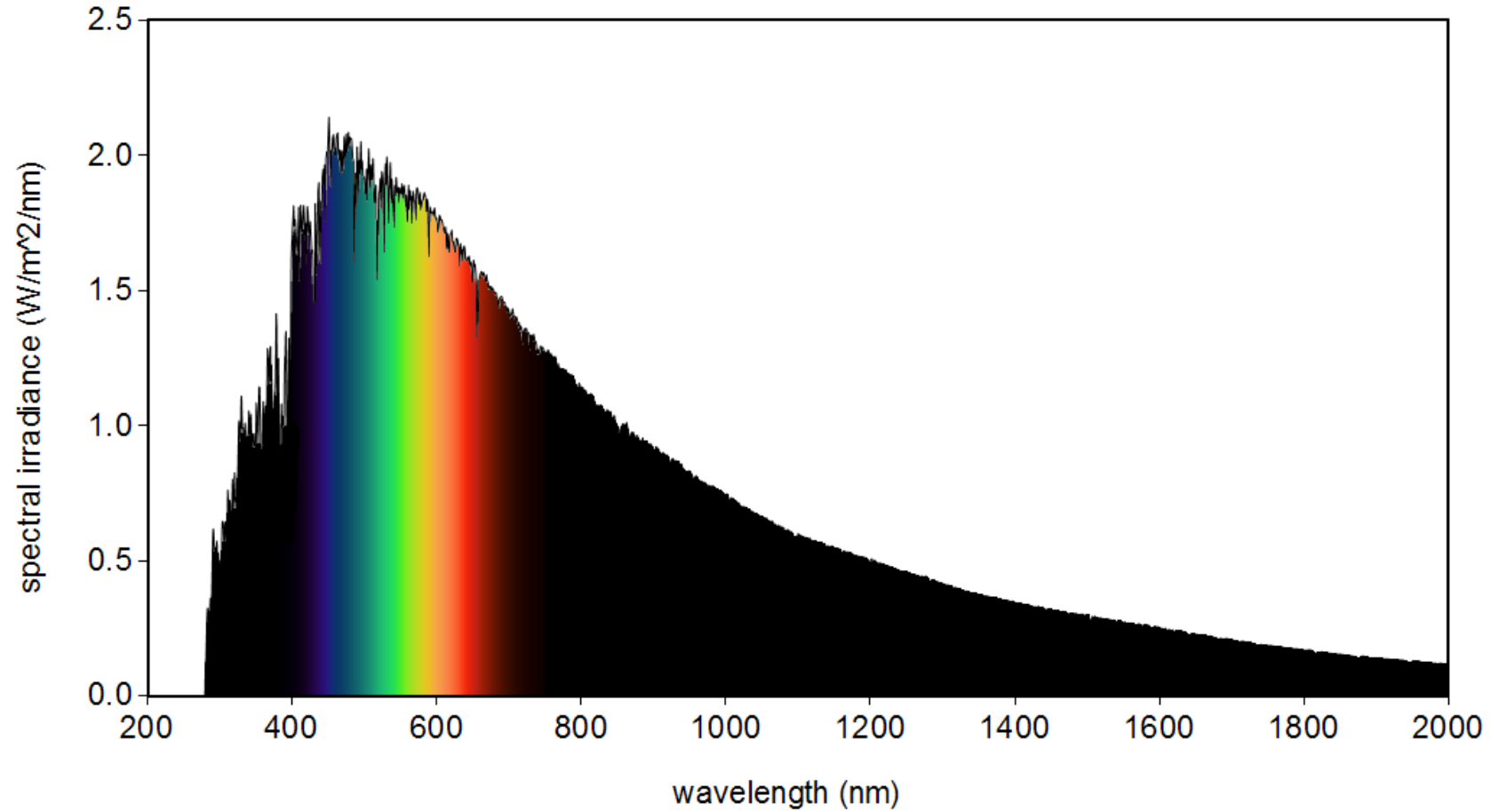


# SPETTRO DELLE ONDE

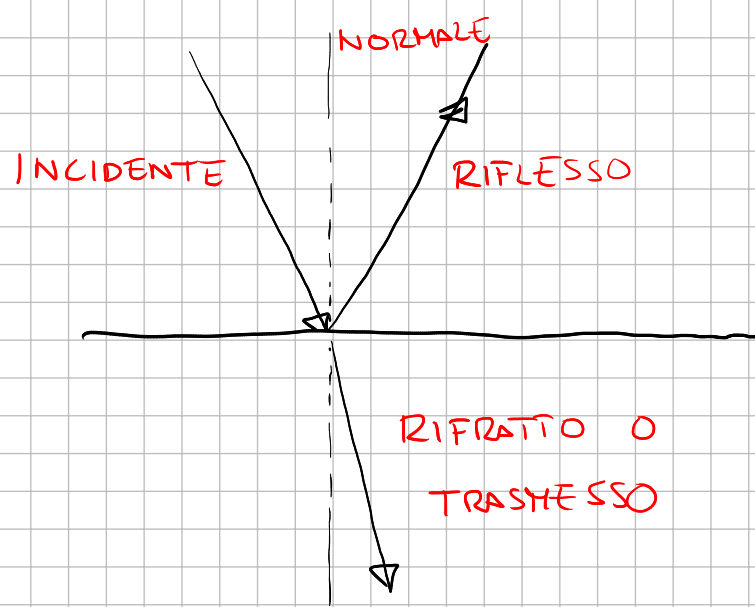


# SPETTRO DEL SOLE

Sunlight spectrum in space as a function of wavelength



## LUCE E INDICE DI RIFRAZIONE



normale + incidente definiscono un piano d'incidenza

riflessi + rifratti giacciono su quel piano

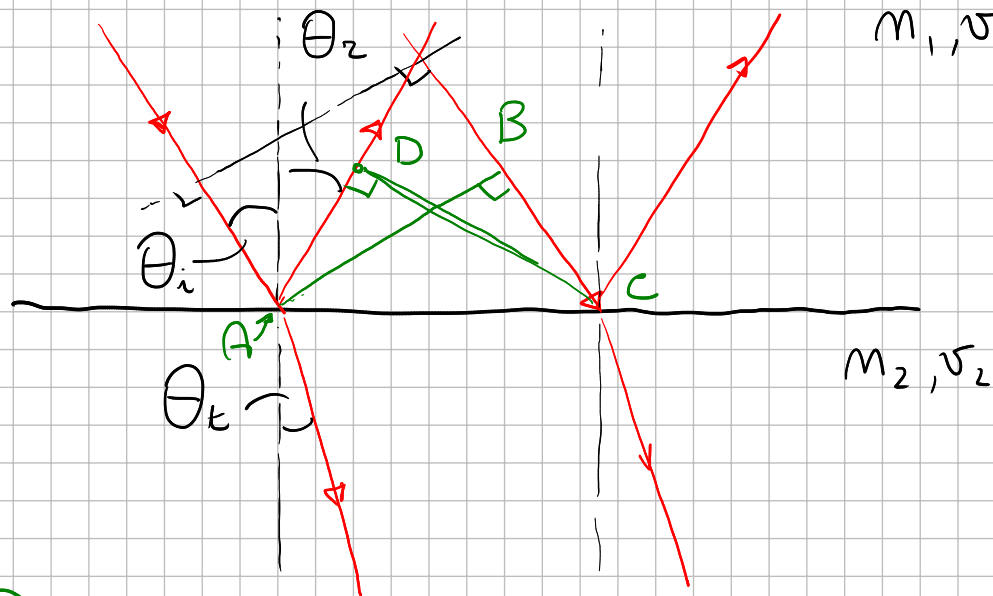
in un mezzo la luce si muove con  $v < c$ ,  $n \equiv \frac{c}{v}$  indice di RIFRAZIONE

$n \geq 1$ ,  $n_{\text{ARIA}} \approx 1$ ,  $n_{\text{ACQUA}} \approx 1.33$ ,  $n_{\text{VETRO}} \approx 1.5 \div 2$

$\omega = k v \Rightarrow v = \frac{\omega}{k} = \frac{\lambda}{T} \rightarrow$  si trova sperimentalmente che

$\lambda = \frac{\lambda_0}{n}$   $\Rightarrow k = n k_0$  perché  $\omega = \omega_0$  [  $\nu = \frac{\omega}{2\pi}$  FREQUENZA DELL'ONDA ]  
 $\nu = \nu_0$

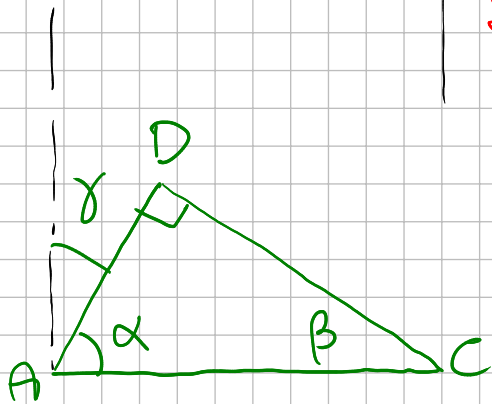
# RIFLESSIONE E RIFRAZIONE



$$\overline{BC} = \overline{AD}$$

guardiamo  $\widehat{ABC}$  e  $\widehat{ADC}$

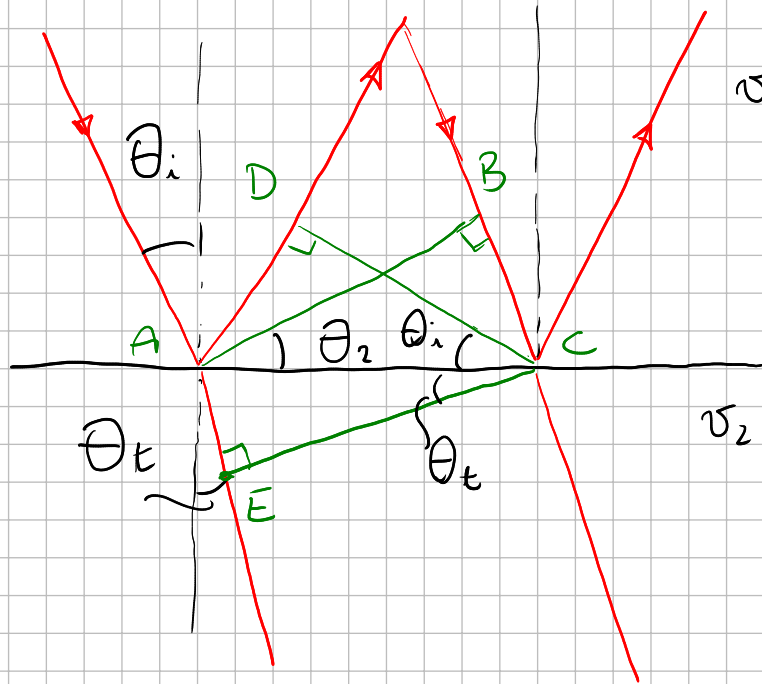
- hanno la stessa ipotenusa,  $\overline{AC}$
  - hanno un angolo retto
  - hanno un cateto uguale,  $\overline{BC} = \overline{AD}$
- $\Rightarrow$  seni uguali  $\rightarrow$  angoli uguali



$$\alpha + \beta = \frac{\pi}{2} \quad \Rightarrow \quad \beta = \frac{\pi}{2} - \alpha$$

$$\gamma + \alpha = \frac{\pi}{2} \quad \Rightarrow \quad \gamma = \frac{\pi}{2} - \alpha = \beta = \theta_2 = \theta_r$$

# RIFRAZIONE



$v_1$

$\overline{BC} \neq \overline{AE}$  perché  $v_1 \neq v_2$

$\widehat{ABC}$  e  $\widehat{AEC}$  hanno  $\overline{AC}$  in comune

$$\left\{ \begin{array}{l} \overline{BC} = \overline{AC} \sin \theta_2 = v_1 \Delta t \\ \overline{AE} = \overline{AC} \sin \theta_t = v_2 \Delta t \end{array} \right. \Rightarrow$$

$$\overline{AC} = \frac{v_1 \Delta t}{\sin \theta_2} = \frac{v_2 \Delta t}{\sin \theta_t} \Rightarrow$$

$$\frac{v_1}{\sin \theta_2} = \frac{v_2}{\sin \theta_t}, \text{ per} \theta_1 = \frac{c}{v_1}, \theta_2 = \frac{c}{v_2} \Rightarrow v_1 = \frac{c}{n_1}, v_2 = \frac{c}{n_2} \Rightarrow$$

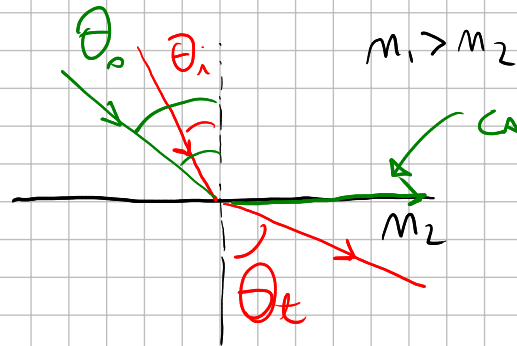
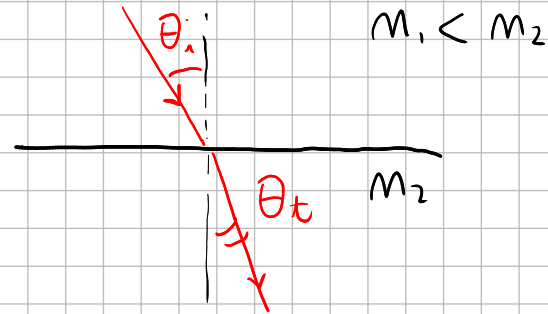
$$n_1 \sin \theta_2 = n_1 \sin \theta_i = n_2 \sin \theta_t \text{ LEGGE DI SNELL}$$

# LEGGÈ DI SNELL

$$\sin \theta_t = \frac{n_1}{n_2} \sin \theta_i$$

$$\sin \theta_t < \sin \theta_i \text{ se } n_2 > n_1$$

$$\sin \theta_t > \sin \theta_i \text{ se } n_1 > n_2$$



CASO LIMITE

$$1 = \frac{n_1}{n_2} \sin \theta_0$$

$\Leftrightarrow$

$$\sin \theta_0 = \frac{n_2}{n_1}$$

# DISPERSIONE DELLA LUCE

